

The effect of single-particle space-momentum angle distribution on two-pion HBT correlation in relativistic heavy-ion collisions using a multiphase transport model

Hang Yang(杨航) Qichun Feng(冯启春) Yanyv Ren(任延宇) Jingbo Zhang(张景波)[†] Lei Huo(霍雷)

School of Physics, Harbin Institute of Technology, Harbin 150001, China

Abstract: Using the string melting version of a multiphase transport (AMPT) model, we analyze the transverse momentum dependence of the HBT radius R_s and the single-pion angle distribution on the transverse plane in central Au+Au collisions at $\sqrt{s_{NN}}=19.6, 27, 39, 62.4,$ and 200 GeV. Additionally, a numerical connection between these two phenomena is established with a series of functions. We can estimate the single-pion angle distribution on the transverse plane from HBT analysis.

Keywords: HBT radii, transverse momentum dependence, space-momentum angle distribution, AMPT

DOI: 10.1088/1674-1137/ac1856

I. INTRODUCTION

The two-pion intensity interferometry method, also called the Hanbury-Brown Twiss (HBT) method, was first performed by Hanbury, Brown, and Twiss to measure the angular diameter of stars in the 1950s [1]. Later, G. Goldhaber, S. Goldhaber, W. Lee, and A. Pais extended this method in particle physics to study the angular distribution of identical pion pairs in $\bar{p}+p$ collisions [2]. Since then, two-pion interferometry has been widely used in high-energy heavy-ion collisions.

In the Relativistic Heavy Ion Collider, a new state of matter, known as a quark-gluon-plasma (QGP), has been found [3]. QGP is strongly interacting partonic matter formed by deconfined quarks and gluons under extreme temperature and energy density. This state is similar to the early time of the universe after the big bang [4]. Furthermore, there is a critical end point (CEP) at the boundaries between the QGP and hadronic gas in the QCD phase diagram [5]. The RHIC collaboration is searching for this CEP using the beam energy scan (BES) program [6]. The HBT method is a useful tool in these investigations, as it can provide the space-time and dynamic information of the freeze-out state of the source [7, 8]. Near this CEP, the abnormal dynamic characteristics of the critical system will lead to a sudden change in the transport coefficients, meaning the HBT method can also be used to estimate the CEP [9, 10].

Many collaborations have already presented some HBT results for different collisions and energies [11-14]. When the HBT radii for the same collision energy were

analyzed, the radii decreased with increasing transverse momentum of the pion pairs. This phenomenon, known as the transverse momentum dependence of HBT radii, is attributed to the space-momentum correlation [15, 16], which is caused by the collective flow [17]. We can obtain more information about the collective flow by studying the transverse momentum dependence of the HBT radii, making it important to study this space-momentum correlation. Hence, we introduce the single-particle space-momentum angle distribution to describe this space-momentum correlation. When the particles freeze out, there will be a finite angle between the radius vector and the momentum vector. Figure 1 displays the diagram of this angle and its projection angle $\Delta\theta$ on the transverse plane. Here, the single-particle space-momentum angle distribution is the quantification of the space-momentum correlation [18]. As this paper focuses on the transverse plane, we only employ the angle $\Delta\theta$.

In this study, a multiphase transport (AMPT) model is used to produce particles and calculate the HBT radius. The AMPT model is a hybrid model, which describes relativistic heavy ion collisions [19], and has already been widely used for HBT analysis [20-22]. We will also focus on how the single-particle space-momentum angle $\Delta\theta$ distribution changes with the transverse momentum dependence on the transverse plane. The HBT radius R_s is directly related to the transverse size of the pion source [23], so we attempt to deduce a numerical connection between the $\Delta\theta$ distribution and the transverse momentum dependence of the HBT radius R_s .

This paper is structured as follows. Sec. II briefly in-

Received 30 March 2021; Accepted 28 July 2021; Published online 3 September 2021

[†] E-mail: jinux@hit.edu.cn

©2021 Chinese Physical Society and the Institute of High Energy Physics of the Chinese Academy of Sciences and the Institute of Modern Physics of the Chinese Academy of Sciences and IOP Publishing Ltd

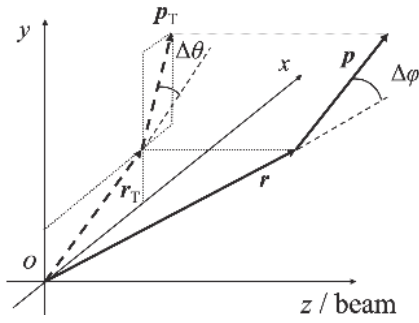


Fig. 1. Diagram of $\Delta\varphi$ and $\Delta\theta$. $\Delta\varphi$ is the angle between r and p , and $\Delta\theta$ is the angle between r_T and p_T , at the freeze-out time. The origin is the center of the source.

roduces the AMPT model and the method used to calculate the HBT radii. In Sec. III, we calculate the HBT radii for pions at different collision energies. In Sec. IV, a numerical connection is constructed between the $\Delta\theta$ angle distribution and the transverse momentum dependence of R_s . Finally, we summarize our conclusions in Sec. V.

II. AMPT MODEL AND METHODOLOGY

The AMPT model is a hybrid model with initial particle distributions generated using the heavy ion jet interaction generator (HIJING) model. The AMPT model consists of two versions, the default AMPT model and the string melting AMPT model. The version employed here is the string melting AMPT model, which provides a good description of the two-pion correlation function [19, 24]. The string melting AMPT model consists of four main components: the initial conditions, partonic interactions, conversion from the partonic to hadronic matter, and hadronic interactions. It can provide better pion phase space information than our previous work, as well as a better $\Delta\theta$ distribution.

The Correlation After Burner (CRAB) code is used to calculate the two-pion correlation functions [25]. The code is based on the formula

$$C(\mathbf{q}, \mathbf{K}) = 1 + \frac{\int d^4x_1 d^4x_2 S_1(x_1, \mathbf{p}_2) S_2(x_2, \mathbf{p}_2) |\psi_{\text{rel}}|^2}{\int d^4x_1 d^4x_2 S_1(x_1, \mathbf{p}_2) S_2(x_2, \mathbf{p}_2)}, \quad (1)$$

where $\mathbf{q} = \mathbf{p}_1 - \mathbf{p}_2$, $\mathbf{K} = (\mathbf{p}_1 + \mathbf{p}_2)/2$, ψ_{rel} is the two particle wave function, and $S(x, \mathbf{p})$ is the emission function of the single particle, which describes the probability of emitting a particle with momentum \mathbf{p} at space-time point x . In further discussion, we neglect the Coulomb interaction and strong interactions between pions and focus on the mid-rapidity range ($-0.5 < \eta < 0.5$).

The HBT three-dimensional correlation function can be written as [26]

$$C(\mathbf{q}, \mathbf{K}) = 1 + \lambda \exp[-q_0^2 R_0^2(\mathbf{K}) - q_s^2 R_s^2(\mathbf{K}) - q_l^2 R_l^2(\mathbf{K})], \quad (2)$$

where λ is the coherence parameter, and R_0 , R_s , and R_l are the HBT radii in the "out-side-long" coordinate system, where o, s, and l are the directions shown in Fig. 2. The longitudinal direction is along the beam direction, while the transverse plane is perpendicular to the longitudinal direction. In the transverse plane, the momentum direction of particle pairs is the outward direction. The direction perpendicular to the outward direction is referred to as the sideward direction. The HBT radii can be calculated using Eq. (2) to fit the HBT correlation function generated by the CRAB code.

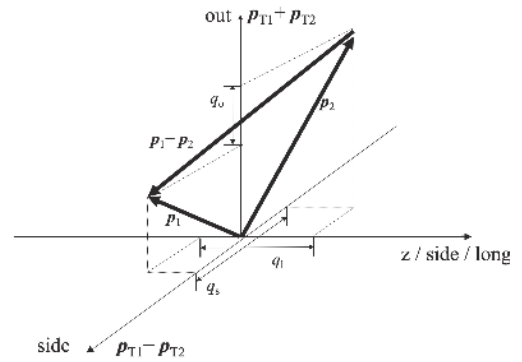


Fig. 2. Diagram of the "out-side-long" (o-s-l) coordinate system.

III. TRANSVERSE MOMENTUM DEPENDENCE OF THE HBT RADIUS

In HBT analysis, R_s is directly related to the transverse size of the emission source, while R_0 and R_l are influenced by the source lifetime and the velocity of the particles [16]. Therefore, R_s is more important in this study, so we focus only on the transverse momentum dependence of R_s .

We employ the melting AMPT model to generate central Au+Au collision events at $\sqrt{s_{NN}} = 19.6, 27, 39, 62.4, 200$ GeV, with these energies chosen from the BES energies and the impact parameter as 0 fm. The Coulomb interaction and strong interactions are neglected, meaning the three types of pions can be treated as the same. The correlation function of the pions is calculated for different transverse momenta by the Crab code. An example of the correlation function of pions is shown in Fig. 3.

Next, we employ Eq. (2) to fit the HBT correlation functions and obtain values of R_s for different transverse momenta at different collision energies, as shown in Fig. 4.

For all collision energies, R_s decreases with increasing transverse momentum of the pairs, and they also

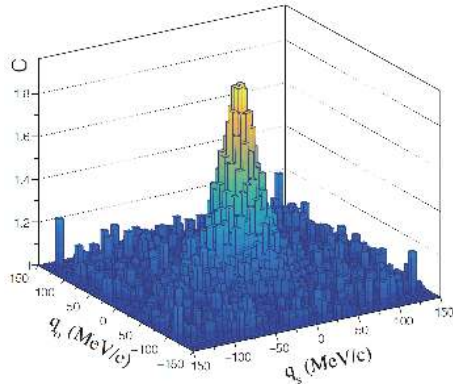


Fig. 3. (color online) Correlation function in the q_0 and q_s directions of the central Au+Au collisions at $\sqrt{s_{NN}} = 200$ MeV for the AMPT model. The K_T range is 175-225 MeV/c, and the q_l range is 0-6 MeV/c.

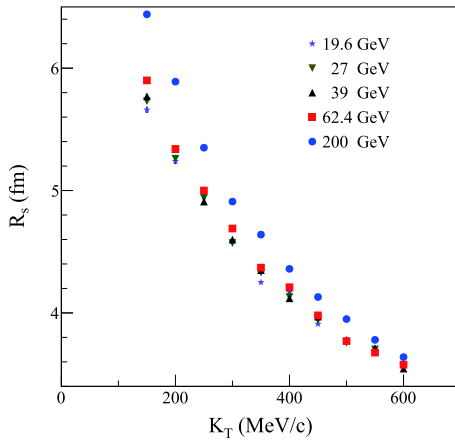


Fig. 4. (color online) Transverse momentum dependence of R_s in AMPT model.

demonstrate different strengths of K_T dependence of R_s . The values of R_s with higher K_T are more similar. We can fit the HBT radii using

$$R = aK_T^b, \quad (3)$$

where parameter a is a common constant, and parameter b reflects how the HBT radii change with K_T ; it describes the strength of the K_T dependence of HBT radii. The collision energy dependence of parameter b is shown in Fig. 5.

In Fig. 5, with increasing collision energy, the values of parameter b for R_s also increase (i.e., a larger $|b|$), indicating a stronger K_T dependence. In our previous study, we found that HBT radii can be influenced by the $\cos(\Delta\theta)$ distribution, and the $\cos(\Delta\theta)$ distribution changes with the pair momentum K_T , leading to changes in the HBT radius R_s and the b value. Thus, we only focus on the dependence of the $\cos(\Delta\theta)$ distribution on the parameter b .

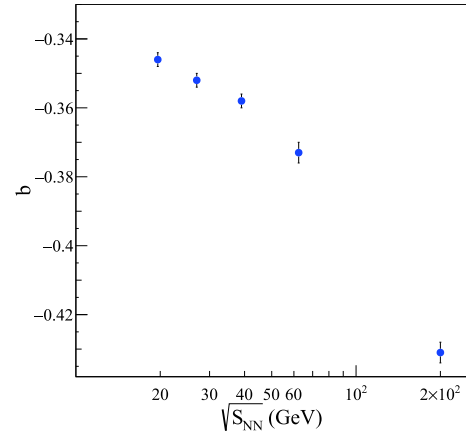


Fig. 5. (color online) Collision energy dependence of parameter b in the AMPT model.

IV. TRANSVERSE SPACE-MOMENTUM ANGLE DISTRIBUTION

When the particles freeze-out, there is an angle between the radius vector and the momentum vector called the single-particle space-momentum angle, defined by $\Delta\theta$ in the transverse momentum plane. We generally calculate the cosine value of this angle. There is substantial particle freeze-out from the source, so we calculate $\cos(\Delta\theta)$ values for all particles to obtain the $\cos(\Delta\theta)$ distribution. This distribution can directly detail the transverse momentum K_T dependence of R_s [18]. We use the pions, which are generated from the AMPT model and used to calculate the HBT correlation function, to obtain the $\cos(\Delta\theta)$ distribution. Pions with random \mathbf{p}_T and \mathbf{r}_T are generated and used to calculate the random $\cos(\Delta\theta)$ distribution. Then, we use the distribution of $\cos(\Delta\theta)$ for different K_T regions, which are calculated from the AMPT model and divided by the random $\cos(\Delta\theta)$ distribution, to obtain the normalized $\cos(\Delta\theta)$ distribution, as shown in Fig. 6.

In Fig. 6, with increasing transverse momentum K_T and collision energies, the normalized $\cos(\Delta\theta)$ distribution becomes closer to $\cos(\Delta\theta) = 1$. This phenomenon indicates that with the increase of transverse momentum K_T and collision energy of the two pions, the momentum direction of a single-pion is closer to the space direction. Additionally, the transverse momentum dependence of HBT can be explained with the transverse momentum dependence of the normalized single-particle space-momentum angle $\Delta\theta$ distribution. In our previous studies, we employed an exponential function to fit the normalized $\Delta\theta$ distribution in a simple cylinder source. While the AMPT model is a hybrid model, the density of the source and the collective flow of the AMPT model is quite different from the cylinder source. Therefore, we modified our fit function to a double exponential to obtain better fitting results. Hence, the parameters set in this paper are

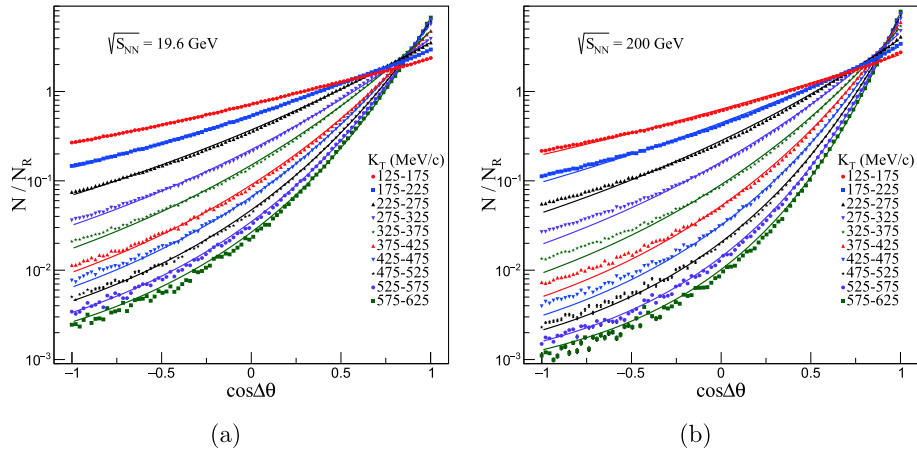


Fig. 6. (color online) Normalized $\cos(\Delta\theta)$ distribution for melting AMPT model.

different from our previous work. The fit functions of the fit lines are

$$f = 0.0005 \exp\left\{c_1 \exp\left[c_2 \cos(\Delta\theta)\right]\right\}, \quad (4)$$

where c_1 and c_2 are fit parameters, and the value 0.0005 was chosen by us to obtain good fitting results. The value of c_1 is influenced by the proportion of the pions whose freeze-out direction is perpendicular to the radius direction, and c_2 is influenced by the strength of the curve approach pion. They are both determined by the normalized $\cos(\Delta\theta)$ distribution and are influenced by the collective flow. The velocity of the flow in different K_T regions is shown in Fig. 7.

In Fig. 7, the pions with higher transverse momentum are emitted by the larger collective flow, with the flow typically strong enough that the direction of freeze-out pions is limited to the flow direction, which means that $\cos(\Delta\theta)$ approaches one. This results in the behavior of the normalized $\cos(\Delta\theta)$ distribution in Fig. 6 and is reflected in the changing values of c_1 and c_2 with the collect-

ive flow. These values are shown in Fig. 8. The tendency is higher for larger K_T and larger collision energy.

Then, we discuss the normalized $\cos(\Delta\theta)$ distribution changing with K_T . We use functions to describe how c_1 and c_2 change with K_T , seen as the fit lines in Fig. 9. These functions are

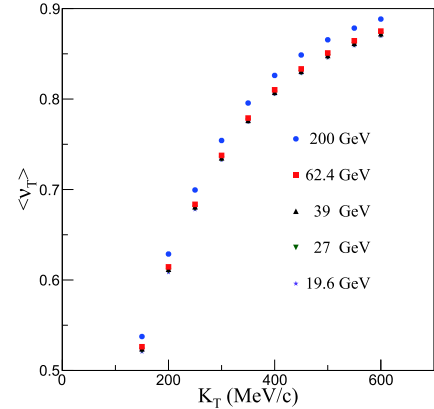


Fig. 7. (color online) Velocity of the flow for melting AMPT model.

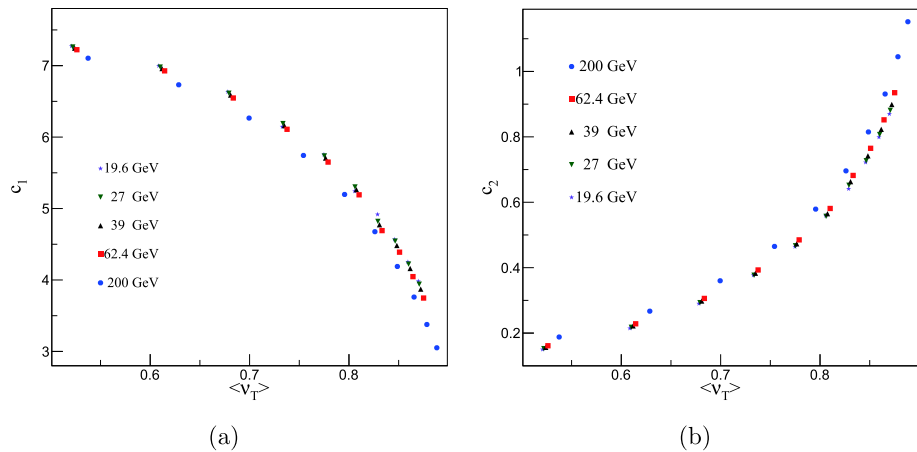


Fig. 8. (color online) Flow velocity dependence of fit parameters for c_1 and c_2 .

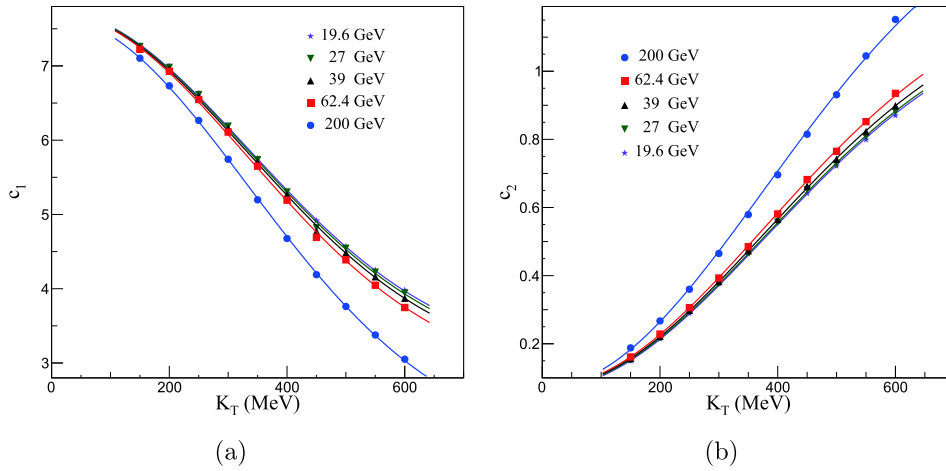


Fig. 9. (color online) Transverse momentum dependence of the fit parameters for c_1 and c_2 .

$$c_1 = k_1 \exp \left[-4.5 \times \left(\frac{K_T}{1000} \right)^2 \right] + j_1, \quad (5)$$

$$c_2 = k_2 \exp \left[-3.5 \times \left(\frac{K_T}{1000} \right)^2 \right] + j_2, \quad (6)$$

where k and j are fit parameters; the values -4.5 and -3.5 were chosen by us to obtain a good fit. The parameters k_1 and j_1 are determined by the changes to c_1 due to K_T . As $K_T \rightarrow +\infty$, the proportion of the pions with $\cos(\Delta\theta) = 0$ is at its lowest, affecting the j_1 values. When $K_T = 0$, this proportion is the highest, with the changing range of this proportion affecting the k_1 values. The parameters k_2 and j_2 are determined by the changes to c_2 with K_T , and when $K_T \rightarrow +\infty$, there is the greatest strength of the normalized $\cos(\Delta\theta)$ distribution approaching one, affecting the j_2 values. When $K_T = 0$, this strength is the lowest, so the changing range of the strength affects the k_2 values. In Fig. 9, at low collision energies, particularly for $\sqrt{s_{NN}} = 19.6, 27,$ and 39 GeV, the collision energies are very similar, as is the strength of collective flow. This results in the single-pion space-momentum angle having a similar distribution within the same transverse momentum region, meaning the c_1 and c_2 values are also similar.

In order to discuss how the normalized $\cos(\Delta\theta)$ distribution changes with the strength of the K_T dependence of R_s , we plot the parameter b , obtained from the transverse momentum dependence of R_s , as a function of the parameters k and j , shown in Fig. 10. The changing pattern is carried by the fit parameters. The pattern indicates that parameter b has an extremum, so the effect of the single-pion space-momentum angle distribution on the HBT radii will be similar at higher energies, and the transverse momentum dependence of the HBT radii will also be similar at higher collision energies. This minimum value was set to -0.5 , and the numerical relationship based on the fit results was then constructed. The red lines depict

the fits, and the fit functions are

$$b(k_1) = \mu_{11}|k_1|^{\mu_{12}} - 0.5, \quad (7)$$

$$b(j_1) = \nu_{11}|j_1|^{\nu_{12}} - 0.5, \quad (8)$$

$$b(k_2) = \mu_{21}|k_2|^{\mu_{22}} - 0.5, \quad (9)$$

$$b(j_2) = \nu_{21}|j_2|^{\nu_{22}} - 0.5. \quad (10)$$

where μ and ν are fit parameters. These parameters are influenced by the change in the normalized $\cos(\Delta\theta)$ distribution with the strength of the K_T dependence of R_s . The fit parameter values are shown in Table 1.

In short, the parameters c_1 and c_2 are determined by the normalized $\cos(\Delta\theta)$ distribution. The proportion of pions whose $\cos(\Delta\theta) = 0$ determines the c_1 values, and the strength of the normalized $\cos(\Delta\theta)$ distribution approaching one determines the c_2 values. The parameters k and j are determined by the distribution changing with pair momentum K_T , with k_1 and j_1 related to c_1 . k_1 is influenced by the changing range of the proportion of pions with $\cos(\Delta\theta) = 0$, and j_1 is influenced by the lowest proportion. k_2 and j_2 are also related to c_2 . k_2 is influenced by the changing range of the strength of the normalized $\cos(\Delta\theta)$ distribution approaching one, and j_2 is influenced by its highest strength. Parameter b describes the

Table 1. Fit results of $b(k)$ and $b(j)$.

	c_1		c_2
μ_{11}	60 ± 20	μ_{21}	0.219 ± 0.005
μ_{12}	-3.9 ± 0.2	μ_{22}	-3.0 ± 0.2
ν_{11}	0.023 ± 0.002	ν_{21}	0.265 ± 0.009
ν_{12}	1.7 ± 0.1	ν_{22}	-3.1 ± 0.2

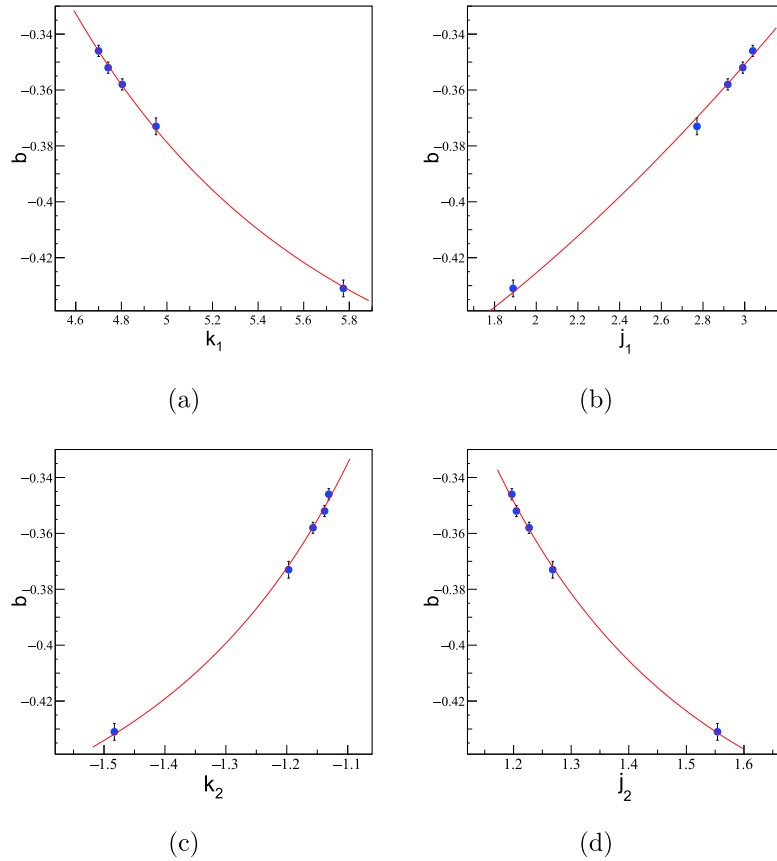


Fig. 10. (color online) Fit parameters in AMPT model. Parameter b is determined from the HBT radii fit function $R = ap_T^b$, and j_1 , j_2 and k_1 , k_2 are based on the fit function $c_1 = k_1 \exp\left[-4.5 \times \left(\frac{K_T}{1000}\right)^2\right] + j_1$ and $c_2 = k_2 \exp\left[-3.5 \times \left(\frac{K_T}{1000}\right)^2\right] + j_2$, where c_1 and c_2 are obtained from the normalized space-momentum angle distribution function $f = 0.0005 \exp\{c_1 \exp[c_2 \cos(\Delta\theta)]\}$. The red lines are fits.

strength of the transverse momentum dependence of R_s , and the parameters μ and ν describe how k and j change with b , which is the connection between the single-pion space-momentum $\cos(\Delta\theta)$ distribution and the transverse momentum dependence of the HBT radius R_s . With the AMPT model, a numerical connection has been established. When we obtain a series of R_s data in different K_T regions, we can estimate the $\cos(\Delta\theta)$ distribution as a function of K_T . With further research, the values chosen can also be improved with more accurate fits.

V. CONCLUSIONS

The transverse momentum dependence of HBT radii is related to the space-momentum correlation, and the single-particle space-momentum angle $\cos(\Delta\theta)$ distribution was used to quantify this correlation. Thus, the transverse momentum dependence of HBT radii can be explained by the transverse momentum dependence of the single-particle space-momentum angle $\cos(\Delta\theta)$ distribution. With the string melting AMPT model, we calculated the transverse momentum of the HBT radius R_s for

several collision energies, with the results suggesting that, with increasing collision energies, the decrease in the HBT radii with the pair momentum is more prominent. Then, we calculated the single-particle space-momentum angle $\cos(\Delta\theta)$ distribution for each pair momentum region and collision energy. The results indicate that, with increasing collision energies and pair momentum, the freeze-out directions of the pions are closer to the radius direction. We use several fit functions to establish a connection between these distributions and the R_s . With this connection, we can describe how the space-momentum correlation changes with increasing collision energies and also obtain information about the final stage of the Au+Au collision at the time the freeze-out occurs. Near the CEP, anomalies can occur in a wide variety of dynamic properties, which can affect the single-particle space-momentum correlation. It is also possible to see some non-monotonic behaviors through HBT analysis. With further research, HBT analysis may help us to learn more about the freeze-out stage of the collisions, and the critical effect of single-particle space-momentum correlations.

References

- [1] R. Hanbury Brown and R. Q. Twiss, *Nature* **178**, 1046-1048 (1956)
- [2] G. Goldhaber, S. Goldhaber, and W. Lee, *Phys. Rev.* **120**(1), 300-312 (1960)
- [3] B. Back, M. Baker, M. Ballintijn *et al.*, *Nucl. Phys. A* **757**, 28-101 (2005)
- [4] H. Satz, *Nucl. Phys. Proc. Suppl.* **94**, 204-218 (2001)
- [5] M. Stephanov, K. Rajagopal, and E. Shuryak, *Phys. Rev. Lett.* **81**(22), 4816-4819 (1998)
- [6] M. M. Aggarwal *et al.* (STAR), 2010 (*Preprint* 1007.2613)
- [7] U. Heinz and B. V. Jacak, *Annu. Rev. Nucl. Part. Sci.* **49**, 529-579 (1999)
- [8] U. A. Wiedemann and U. Heinz, 1999 *Phys. Rep.* 319 145 - 230 ISSN 0370-1573
- [9] R. A. Lacey, *Phys. Rev. Lett.* **114**(14), 142301 (2015)
- [10] R. A. Lacey, N. N. Ajitanand, J. M. Alexander *et al.*, 2007(*Preprint* 0708.3512)
- [11] J. Adams, M. M. Aggarwal, Z. Ahammed *et al.* (STAR Collaboration), *Phys. Rev. C* **71**(4), 044906 (2005)
- [12] S. Kniege and (for the NA49 Collaboration), *J. Phys. G-Nucl. Part. Phys.* **30**, S1073-S1077 (2004)
- [13] K. Aamodt, A. Abrahantes, D. Quintana, Adamová *et al.* (ALICE Collaboration), *Phys. Rev. D* **84**(11), 112004 (2011)
- [14] L. Adamczyk, J. K. Adkins, G. Agakishiev *et al.* (STAR Collaboration), *Phys. Rev. C* **92**(1), 014904 (2015)
- [15] M. A. Lisa, S. Pratt, R. Soltz *et al.*, *Annu. Rev. Nucl. Part. Sci.* **55**, 357-402 (2005)
- [16] U. A. Wiedemann, P. Scotto, and U. Heinz, *Phys. Rev. C* **53**(2), 918-931 (1996)
- [17] S. Pratt, *Phys. Rev. Lett.* **53**(13), 1219-1221 (1984)
- [18] H. Yang, Q. Feng, Y. Ren *et al.*, *Chin. Phys. C* **44**, 054105 (2020)
- [19] Z. W. Lin, C. M. Ko, B. A. Li *et al.*, *Phys. Rev. C* **72**(6), 064901 (2005)
- [20] Y. Zhang, J. Zhang, J. Liu *et al.*, *Phys. Rev. C* **92**(1), 014909 (2015)
- [21] Y. Zhang, J. Zhang, T. Chen *et al.*, *Phys. Rev. C* **96**(4), 044914 (2017)
- [22] L. Q. Shan, F. J. Wu, J. L. Liu *et al.*, *J. Phys. G-Nucl. Part. Phys.* **36**, 115102 (2009)
- [23] K. Aamodt, A. A. Quintana, D. Adamová *et al.*, *Phys. Lett. B* **696**, 328-337 (2011)
- [24] Z. w. Lin, C. M. Ko, and S. Pal, *Phys. Rev. Lett.* **89**(15), 152301 (2002)
- [25] S. Pratt, 2006 Crab version 3.0 <https://web.pa.msu.edu/people/pratts/freecodes/crab/home.html>
- [26] G. Alexander, *Rep. Prog. Phys.* **66**, 481-522 (2003)